

# Field Manual of Fine-Tuning

## Chapter Notes: Particle Physics

*Author:* Giulia Zanderighi  
*Interviewer:* Rafael Alves Batista  
*Date of Interviews:* 31 May 2016  
*Date Deliverable:* ?? Mar 2016  
*Audio Recorded?* Yes  
*Video Recorded?* No  
*Workshop Attendance:* TBA

### Physics Covered

Particle Physics

### Previous Chapter:

Unknown, possibly fine-structure constants

### Following Chapter:

Unknown, likely nucleosynthesis

### Types of Fine-Tuning:

X small, for fundamental constants X such that process still happens. Fine-tuning of coefficients to account for renormalisability.

$A/B$  (dimensionless) too large/small, where  $A$  and  $B$  are ratio of quantities such as masses.

### Directly explored Fine-Tuned Parameters:

- $v_H$ : vacuum expectation value of the Higgs field
- $m_H$ : mass of the Higgs boson
- $m_q$ : mass of quarks
- $m_l$ : mass of charged leptons
- $m_\nu$ : mass of neutrinos

### Indirectly explored Fine-Tuned Parameters:

- $m_{Z^0, W^\pm}$ : mass of electroweak gauge bosons
- $m_g$ : mass of the gluon
- $\theta_W$ : Weinberg angle
- $\Delta m_{\nu, ij}$ : mass hierarchy in neutrino sector
- $g_w$ : electroweak coupling constant
- $g_s$ : strong coupling constant
- $\alpha_e$ : electroweak fine structure constant
- $\alpha_s$ : strong fine structure constant
- $\Delta m_{p, n}$ : mass difference between protons and neutrons
- $\Delta m_{u, d}$ : mass difference between up and down quarks.

### Assumed Background:

- quantum field theory

**Outline of Chapter:**

## Introduction: The Standard Model

- There is no way to properly quantify fine-tunings in particle physics.
- Fine-tuning in particle physics is not related to initial conditions, but rather to values of parameters.
- “First principle” explanation for masses and couplings: why is there such a big scale gap between different particle masses.

## Naturalness

- What is naturalness, why is it important and why we expect it?
- Isn't naturalness simply an “aesthetic” argument?
- Why do we adopt naturalness as a valid scientific driver? Historical examples.
- “If parameters do not seem natural, there is usually a reason”, e.g. symmetry, new physics.
- Degeneracy of  $c\Lambda^2 \rightarrow$  why expect  $c \sim 1$  rather than a smaller  $\Lambda$ ?
- Why expect naturalness of parameters instead of naturalness of a function of these parameters.

## Hierarchy problem

- Higgs mass seems closer to electroweak than to Planck scale without an apparent reason.
- Consequences of  $m_H - m_{Pl} \ll m_H - m_{EW}$  ?
- “Little hierarchy problem” and consequences for supersymmetry.

## Strong-CP problem

- What is the strong-CP problem and what are possible solutions?
- Usually two approaches are adopted: either we introduce a new particle (in this case the axion), or we change our theory. People have been opting for the former for decades. At what stage does something become so unnatural/unusual that it is worth introducing some new particle to preserve your symmetry?
- Are there possible fine-tunings in the  $\theta_{CP}$  angle? If  $\theta_{CP}$  were much larger or much smaller it would affect matter-genesis (?).

## Flavours

- Why is  $m_t \sim E_{EW}$ ? Is it just a coincidence?
- Why are the masses of particles in different generations so different? Why don't these masses scale equally?  $\rightarrow$  SM is an effective theory
- Why is our universe made of 1st generation particles?

## Higgs

- Higgs vev ( $v_H$ ) is  $\approx 246$  GeV. This is fixed by the potential. Why the potential has this shape?
- Is  $v_H$  more of a fixed parameter coming from the theory, or is it an initial condition “dialed” from a distribution like the cosmological constant in some models?

## Standard Model - Problems

- The standard Standard Model does not predict neutrino masses.
- The SM is more of an effective theory than a fundamental one.
- There are 28 free parameters in the SM. How many of these are actual free parameters, and how many may be possibly related to others? (e.g.:  $m_t$  and  $m_c$  can be written in terms of  $m_u$ , and so there are two fewer free parameters).
- We would still have  $\sim 10$  free parameters, unless we have unification.
- Terms of  $\mathcal{O}(4)$  should be set to zero to guarantee renormalisability. Higher order terms would converge if coefficients are of order 1. No real physical motivation to expect coefficients to be  $\mathcal{O}(1)$ .

- If the SM is not just an effective theory, but part of a more fundamental physical theory, it should work for the case of curved space-times. But QFT doesn't work with curvature.

#### Others

- Do extra dimension models solve fine-tuning?
- Alleviating fine-tuning with extra dimension models: e.g.: technicolor, brane world, Randall-Sundrum

#### Fine-Tuning Issues

- Mass of top  $m_t$  may be totally different than what it is without affecting anything in the universe, provided that  $m_t \gg m_u$  and  $m_t \ll$  energy scale of symmetry breakings.
- Anthropic consequences for the mass of neutrinos and neutrino mass hierarchy: if, for the sake of argument, we had a universe made of 3rd generation particles instead of 1st, then an inverted mass hierarchy would imply that processes involving neutrino production would carry a smaller fraction of energy (rest mass) than the processes we have now. This might have consequences, for example, for supernovae (huge amounts of energy are emitted as  $\nu$ s).
- Consequences of changing the Higgs vev to another value.
- Even if we had  $m_u = m_d$ , it is not clear that  $m_p = m_n$ .
- Stability of atoms for  $m_e \gg m_n - m_p$  and  $m_e \ll m_n - m_p$ .
- $\pi^0$  as p-n mediator: is there any fine-tuning?
- Possible relation between neutrino mass and supernovae.